

A Derivative Expansion for the Casimir Effect

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Introduction

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- *Based on a time-honored idea by Derjaguin:*
 - B.V. Derjaguin, Koll. Z. 69, N2, 155 (1934);
 - B.V. Derjaguin and I. I. Abrikosova, Sov. Phys. JETP **3**, 819 (1957);
 - B.V. Derjaguin, Sci. Am. **203**, 47 (1960).

Sketch of the PFA

Two surfaces, one flat (L), the other (R) describable by a single Monge patch:

$$L) x_3 = 0$$

$$R) x_3 = \psi(x_1, x_2)$$

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$$\begin{aligned} L) x_3 &= 0 \\ R) x_3 &= \psi(x_1, x_2) \end{aligned}$$

The PFA energy is:

$$E_{PFA} = \int d^2 \mathbf{x}_{\parallel} \mathcal{E}_{pp}[\psi(\mathbf{x}_{\parallel})]$$

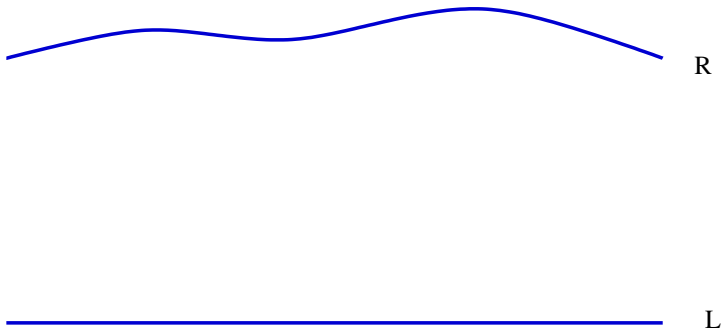
$\mathcal{E}_{pp}(a)$: energy per unit area for two parallel plates.

For a massless real scalar field and two (infinite) parallel Dirichlet surfaces at a distance a :

$$\mathcal{E}_{pp}(a) = -\frac{\pi^2}{1440 a^3},$$

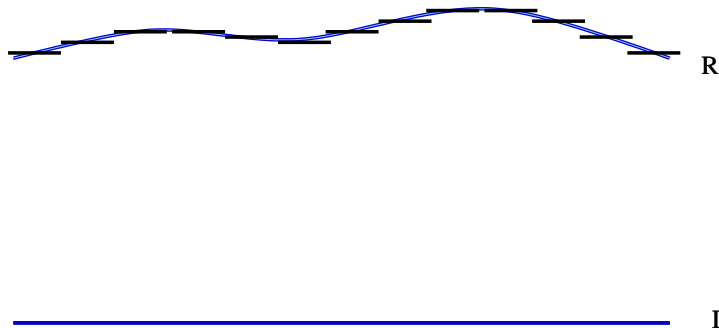
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Each plane contributes: $d^2\mathbf{x}_{\parallel} \mathcal{E}_{pp}[\psi(\mathbf{x}_{\parallel})]$

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- It is quite clear that a way to improve the PFA, while trying to keep its simplicity, would be very welcome.
- In particular, we would like to be able to assess the (in)accuracy of the PFA, without having to solve the exact problem.
- In other words, going beyond the PFA at the lowest possible cost.

Derivative expansion (DE) approach

I will talk about results presented in:

F.D. Lombardo, F.D.Mazzitelli and C. D. F.; Phys.Rev. D84 (2011) 105031;
Phys. Rev. D86 (2012) 045021;
Phys. Rev. D85 (2012) 125037

Interesting generalizations and extensions in:

G. Bimonte, T. Emig, R. L. Jaffe, M. Kardar. Europhys.Lett.,97, 50001 (2012);
G. Bimonte, T. Emig, M. Kardar. Appl.Phys.Lett.,100,074110 (2012).

Motivation for a DE (single Monge patch)

- The Casimir energy E is a **functional** of the *shape* (i.e. ψ) of the surfaces. PFA should be the form adopted by the functional for *almost flat* surfaces.

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- The Casimir energy E is a **functional** of the *shape* (i.e. ψ) of the surfaces. PFA should be the form adopted by the functional for *almost flat* surfaces.
- Thus, the expansion we look for should reproduce, to lowest order, the PFA.
- Terms involving **derivatives** of ψ should appear in the corrections to the PFA.

Idea of the DE

In QFT, there is an approach to evaluating the **effective action** that does exactly that:

- C. M. Fraser, Z. Phys. C 28, 101 (1985); I. J. R. Aitchison and C. M. Fraser, Phys. Lett. 146B, 63 (1984); *ibid.* Phys. Rev. D 31, 2605 (1985).
- R. Jackiw, Phys. Rev. D 9, 1686 (1974).
- L. H. Chan, Phys. Rev. Lett. 54, 1222 (1985); *ibid.* 55, 21 (1985).

Basic idea

To perform an expansion of the effective action (due to an integrated fluctuating field) according to the number of derivatives of its arguments (the 'classical' field).

In the Casimir context: classical field = ψ .

Vacuum energy from the effective action

Total vacuum energy E :

$$E = \lim_{T \rightarrow \infty} \left[\frac{1}{T} \Gamma(\psi) \right]$$

where:

$$\Gamma = -\log \frac{\mathcal{Z}[\psi]}{\mathcal{Z}_0}$$

- T : length of the (imaginary) time interval;
- $\mathcal{Z}[\psi]$: vacuum amplitude in the presence of the mirrors;
- \mathcal{Z}_0 : free (no surfaces) vacuum amplitude;
- (E : also the $\beta \rightarrow \infty$ limit of the free energy);
- We shall subtract self-energies.
- ψ : time-independent (this may be imposed at the end of the calculation).

Form of the DE for Γ (2nd order)

Up to order two in derivatives of ψ :

$$\Gamma[\psi] = \int d^3x_{\parallel} \left[V_{\text{eff}}(\psi) + z(\psi)(\partial\psi)^2 \right].$$

$$(x_{\parallel} \equiv (x_0, x_1, x_2))$$

Constraints

1) Dimensionful quantities: ψ and its derivatives; 2) It should reduce to the PFA for sufficiently smooth surfaces; 3) Locality.

Then:

$$\Gamma[\psi] = \int d^3x_{\parallel} \frac{1}{\psi(x_{\parallel})^3} \left[b_0 + b_2 (\partial\psi)^2 \right].$$

b_0 and b_2 : dimensionless numbers that completely determine the approximation. Note the 'universality' of the coefficients, namely, they are the same, regardless of the (smooth) surface considered.

'Constructive' approach

The previous equation may also be obtained in a more constructive way. If one knows $\Gamma(\psi)$ for:

- $\psi(x_{\parallel}) = a + \eta(x_{\parallel})$, $|\eta| \ll a$, to second order in η .
- Second order in derivatives of η , such that:

$$\Gamma(a + \eta) \simeq \int d^3x_{\parallel} \left[V_{\text{eff}}(a) + z(a)(\partial\psi)^2 \right],$$

then, to second order in derivatives of ψ (and no assumption about magnitude of η):

$$\Gamma[\psi] = \int d^3x_{\parallel} \left[V_{\text{eff}}(\psi) + z(\psi)(\partial\psi)^2 \right].$$

Namely: replace a by ψ in the a -dependent coefficients of the expansion.

The proof relies on the resummation of all the orders in η , in the expansion of Γ to second order in derivatives.

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Determination of the b_0, b_2 coefficients

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- More general approach: to determine them from the expansion of Γ for:

$$\psi(x_{\parallel}) = a + \eta(x_{\parallel})$$

with $|\eta| \ll a$ and smooth, to second order in derivatives of η .

- Then b_0 and b_2 may be determined from the knowledge of E up to the second order in η and second order in derivatives. In particular, $b_0 = -\frac{\pi^2}{1440}$ is derived from the pp result (real scalar field).

Effective action (Dirichlet)

We need the effective action for a scalar field with Dirichlet Boundary Conditions.

Based on the functional representation for the CE:

H. Li and M. Kardar, PRA **46**, 6490 (1992); T. Emig, A. Hanke, R. Golestanian, and M. Kardar, PRL **87**, 260402 (2001); *ibidem* Phys. Rev. **A67**, 022114 (2003).

Functional representation of $\mathcal{Z}(\psi)$:

$$\mathcal{Z}(\psi) = \int \mathcal{D}\varphi \delta_L(\varphi) \delta_R(\varphi) e^{-S_0(\varphi)},$$

$\delta_A(\varphi)$, $A = L, R$: functional δ function which imposes D conditions on the respective mirror, S_0 : free Euclidean action:

$$S_0 = \frac{1}{2} \int dx_0 \int d^3x (\partial\varphi)^2.$$

Auxiliary fields

Exponentiate δ -functionals with two auxiliary fields, λ_L and λ_R , functions of $x_{\parallel} \equiv (x_0, x_1, x_2) \equiv (x_0, \mathbf{x}_{\parallel})$, also satisfying periodic boundary conditions in the x_0 coordinate.

Dirichlet case:

$$\delta_L(\varphi) = \int \mathcal{D}\lambda_L e^{i \int d^3 x_{\parallel} \lambda_L(x_{\parallel}) \varphi(x_{\parallel}, 0)}$$

$$\delta_R(\varphi) = \int \mathcal{D}\lambda_R e^{i \int d^3 x_{\parallel} \sqrt{g(x_{\parallel})} \lambda_R(x_{\parallel}) \varphi(x_{\parallel}, \psi(x_{\parallel}))},$$

g is the determinant of $g_{\alpha\beta}$, the induced metric on R :

$$g_{\alpha\beta}(x_{\parallel}) = \delta_{\alpha\beta} + \partial_{\alpha}\psi(x_{\parallel})\partial_{\beta}\psi(x_{\parallel}),$$

Derivation

Using the exponential representations above and integrating over φ ,

$$\mathcal{Z}(\psi) = \mathcal{Z}^{(0)} \int \mathcal{D}\lambda_L \mathcal{D}\lambda_R e^{-\frac{1}{2} \int_{x_{\parallel}, x'_{\parallel}} \lambda_A(x_{\parallel}) \mathbb{T}_{AB}(x_{\parallel}, x'_{\parallel}) \lambda_B(x'_{\parallel})},$$

where a self-energy contribution has been discarded.

The \mathbb{T} Dirichlet kernel matrix

$$T_{LL}(x_{\parallel}, x'_{\parallel}) = \langle x_{\parallel}, 0 | (-\partial^2)^{-1} | x'_{\parallel}, 0 \rangle \quad (1)$$

$$T_{LR}(x_{\parallel}, x'_{\parallel}) = \langle x_{\parallel}, 0 | (-\partial^2)^{-1} | x'_{\parallel}, \psi(x'_{\parallel}) \rangle \quad (2)$$

$$T_{RL}(x_{\parallel}, x'_{\parallel}) = \langle x_{\parallel}, \psi(x_{\parallel}) | (-\partial^2)^{-1} | x'_{\parallel}, 0 \rangle \quad (3)$$

$$T_{RR}(x_{\parallel}, x'_{\parallel}) = \langle x_{\parallel}, \psi(x_{\parallel}) | (-\partial^2)^{-1} | x'_{\parallel}, \psi(x'_{\parallel}) \rangle \quad (4)$$

$$\langle x | (-\partial^2)^{-1} | y \rangle = \int \frac{d^4 k}{(2\pi)^4} \frac{e^{ik \cdot (x-y)}}{k^2} \equiv \Delta(x-y).$$

2nd order expansion

Now, split ψ into two components,

$$\psi(\mathbf{x}_{\parallel}) = a + \eta(\mathbf{x}_{\parallel}) .$$

a : spatial average of ψ and η contains its varying piece.

Expand Γ up to the second order in η , keeping up to the second order term in an expansion in derivatives.

Expanding the matrix \mathbb{T} in powers of η

$$\mathbb{T} = \mathbb{T}^{(0)} + \mathbb{T}^{(1)} + \mathbb{T}^{(2)} + \dots ,$$

we obtain $\Gamma = \Gamma^{(0)} + \Gamma^{(1)} + \Gamma^{(2)} + \dots$, where

$$\Gamma^{(0)} = \frac{1}{2} \text{Tr} \log \mathbb{T}^{(0)} , \quad \Gamma^{(1)} = \frac{1}{2} \text{Tr} \left[(\mathbb{T}^{(0)})^{-1} \mathbb{T}^{(1)} \right] ,$$

$$\Gamma^{(2)} = \frac{1}{2} \text{Tr} \left[(\mathbb{T}^{(0)})^{-1} \mathbb{T}^{(2)} \right] - \frac{1}{4} \text{Tr} \left[(\mathbb{T}^{(0)})^{-1} \mathbb{T}^{(1)} (\mathbb{T}^{(0)})^{-1} \mathbb{T}^{(1)} \right]$$

Zeroth-order term (PFA)

Zeroth-order term may be obtained by replacing first ψ by a constant, a , and then subtracting the contribution corresponding to $a \rightarrow \infty$. This yields,

$$\Gamma^{(0)} = \frac{1}{2} \text{Tr} \log [1 - (T_{LL}^{(0)})^{-1} T_{LR}^{(0)} (T_{RR}^{(0)})^{-1} T_{RL}^{(0)}]$$

$T_{\alpha\beta}^{(0)}$: identical to the ones for two flat parallel mirrors at a distance a apart. After evaluating the trace:

$$V_{\text{eff}}(\psi) = \frac{1}{2} \int \frac{d^3 k_{\parallel}}{(2\pi)^3} \log [1 - e^{-2k_{\parallel} \psi(x_{\parallel})}] ,$$

which yields the PFA approximation to the vacuum energy.

Second order term

(First order term vanishes identically.) For the second order one we have two contributions:

$$\Gamma^{(2)} = \Gamma^{(2,1)} + \Gamma^{(2,2)}$$

$$\Gamma^{(2,1)} = \frac{1}{2} \text{Tr} \left[(\mathbb{T}^{(0)})^{-1} \mathbb{T}^{(2)} \right]$$

and

$$\Gamma^{(2,2)} = -\frac{1}{4} \text{Tr} \left[(\mathbb{T}^{(0)})^{-1} \mathbb{T}^{(1)} (\mathbb{T}^{(0)})^{-1} \mathbb{T}^{(1)} \right],$$

where we have to keep up to two derivatives of η .

In Fourier space, and before expanding to second order in momentum (derivatives):

$$\Gamma^{(2,j)} = \frac{T}{2} \int \frac{d^2 k}{(2\pi)^2} f^{(2,j)}(\mathbf{k}) |\tilde{\eta}(\mathbf{k})|^2$$

($j = 1, 2$), $\mathbf{k} = (k_1, k_2)$, $f^{(2,j)}$ are the kernels:

$$f^{(2,1)}(k) = - \int \frac{d^3 p}{(2\pi)^3} \frac{|p| |p+k|}{1 - e^{-2|p+k|a}}$$

$$f^{(2,2)}(k) = - \int \frac{d^3 p}{(2\pi)^3} \frac{|p| |p+k| e^{-2|p+k|a} (1 + e^{-2|p|a})}{(1 - e^{-2|p|a})(1 - e^{-2|p+k|a})}.$$

We need to subtract an a -independent (self-energy) contribution.
 Putting together the results from the two terms above:

$$\Gamma^{(2)} = \frac{T}{2} \int \frac{d^2 k}{(2\pi)^2} f^{(2)}(\mathbf{k}) |\tilde{\eta}(\mathbf{k})|^2$$

with:

$$f^{(2)}(k) = -2 \int \frac{d^3 p}{(2\pi)^3} \frac{|p| |p+k|}{(1 - e^{-2|p|a})(e^{2|p+k|a} - 1)}$$

where we just need to extract its \mathbf{k}^2 term in a Taylor expansion at zero momentum. Namely $f^{(2)}(\mathbf{k}) \simeq \chi \mathbf{k}^2$, where

$$\begin{aligned} \chi &= \frac{1}{2} \left[\frac{\partial^2 f^{(2)}(k)}{\partial k^2} \right]_{k \rightarrow 0} \\ &= - \int \frac{d^3 p}{(2\pi)^3} \frac{|p|}{(1 - e^{-2|p|a})} \lim_{k \rightarrow 0} \frac{\partial^2}{\partial k^2} \left[\frac{|p+k|}{(e^{2|p+k|a} - 1)} \right]. \end{aligned}$$

The resulting integral may be exactly calculated,

$$\chi = -\frac{\pi^2}{1080 a^3}.$$

Thus,

$$\begin{aligned}\Gamma^{(2)}(a, \eta) &= -\frac{T}{2} \frac{\pi^2}{1080} \int \frac{d^2 k}{(2\pi)^2} \frac{\mathbf{k}^2}{a^3} |\tilde{\eta}(\mathbf{k})|^2 \\ &= -\frac{T}{2} \frac{\pi^2}{1080} \int d^2 \mathbf{x}_{\parallel} \frac{1}{a^3} (\partial_{\alpha} \eta)^2,\end{aligned}$$

where, to obtain the second order contribution in derivatives to the vacuum energy, we need to replace $a \rightarrow \psi$, $\eta \rightarrow \psi$, and cancel the T factor, obtaining:

$$E_{\text{vac}}^{(2)} = \frac{\Gamma^{(2)}(\psi)}{T} = -\frac{1}{2} \frac{\pi^2}{1080} \int d^2 \mathbf{x}_{\parallel} \frac{(\partial_{\alpha} \psi)^2}{\psi^3},$$

(α runs from 1 to 2).

The final result

Putting together the terms, the DE to second order yields:

$$E_{DE} \equiv -\frac{\pi^2}{1440} \int d^2\mathbf{x}_{\parallel} \frac{1}{\psi^3} \left[1 + \frac{2}{3} (\partial_{\alpha}\psi)^2 \right].$$

- Remarkably simple formula.
- First term: PFA for the Casimir energy.
- Second term: contains first non-trivial correction to the PFA.
- A straightforward calculation provides the relative weight between both terms.
- It is local.
- Amounts to using a better local approximation to the energy density for two surface elements.
- Other (perfect) boundary conditions: different constant coefficients.

Sinusoidal corrugations in x_1 direction

$$\psi(x_1) = a + \epsilon \sin \frac{2\pi x_1}{\lambda} .$$

$$E_{DE} = \frac{\pi^2}{1440} \left[\int d^2 \mathbf{x}_{\parallel} \frac{1}{(a + \epsilon \sin \frac{2\pi x_1}{\lambda})^3} \left(1 + \frac{2}{3} \left(\frac{2\pi}{\lambda} \right)^2 \epsilon^2 \cos^2 \frac{2\pi x_1}{\lambda} \right) \right]$$

In this case, the derivative expansion should be reliable when λ is much larger than both ϵ and a . Should therefore be the largest relevant distance in the problem.

Expanding for $\epsilon \ll a$,

$$E_{DE} = \frac{\pi^2 L^2}{1440 a^3} \left[1 + 3 \left(\frac{\epsilon}{a} \right)^2 + \frac{4\pi^2}{3} \left(\frac{\epsilon}{\lambda} \right)^2 \right] .$$

in perfect agreement with the expansion of the result by Emig et al, when expanded for $a \ll \lambda$.

A sphere in front of a plane

Sphere of radius R at a distance d from a plane.

Expect the derivative expansion to work well in the large- R limit.

$$\psi(\rho) = a + R \left(1 - \sqrt{1 - \frac{\rho^2}{R^2}} \right).$$

(hemisphere, $0 \leq \rho \leq R$). DE well defined if we restrict the integrations to the region $0 \leq \rho \leq \rho_M < R$.

The DE integral can be evaluated. Results in a rather long analytic expression $E_{DE}(\rho_M, a, R)$.

Expanded in inverse powers of R^{-1} :

$$E_{DE} \simeq \frac{\pi^3}{1440} \frac{R}{a^2} \left(1 + \frac{1}{3} \frac{a}{R} \right).$$

Up to this order, the result does not depend on ρ_M . In agreement with the asymptotic expansion obtained from the exact formula for this configuration

Cylinder in front of a plane

Cylinder of radius R at a distance a from a plane. Function ψ given by:

$$\psi(x_1) = a + R \left(1 - \sqrt{1 - \frac{x_1^2}{R^2}} \right),$$

with $-x_M < x_1 < x_M < R$ (to cover the part of the cylinder which is closer to the plane). Result:

$$E_{DE} \simeq \frac{\pi^3 L}{1920\sqrt{2}} \frac{R^{1/2}}{a^{5/2}} \left(1 + \frac{7}{36} \frac{a}{R} \right).$$

The result does not depend on x_M . In agreement with the asymptotic expansion obtained from the exact formula.

Higher dimensions and thermal effects

F.D.L, F.D.M. and C.D.F. Phys. Rev. D86 (2012) 045021

Is the DE always well-defined?

What we have found (d spatial dimensions):

- Real scalar field with either D or N boundary conditions on both mirrors (one flat, one gently curved).
- D boundary conditions: NTLO term is quadratic in derivatives ($\forall d$).
- N boundary conditions: the same holds true for $d \neq 2$. $d = 2$, NTLO term becomes nonlocal. Same nonlocality for $d = 3$ ('dimensional reduction') at high temperatures.
- N mirrors in $d = 3$, we find a nonlocal NTLO term for any $T > 0$.

The system

Free energy: $\Gamma_\beta(\psi)$,

Vacuum energy: $E_{\text{vac}}(\psi) = \lim_{\beta \rightarrow \infty} \Gamma_\beta(\psi)$.

$$\Gamma_\beta(\psi) = -\frac{1}{\beta} \log \left[\frac{\mathcal{Z}_\beta(\psi)}{\mathcal{Z}_\beta^{(0)}} \right].$$

Derivative expansion for D

Same idea and approach, but different number of spatial dimensions d . Time coordinate is compact (periodic).

Keeping up to two derivatives, the Casimir free energy can be written as follows:

$$\Gamma_{\beta}(\psi) = \int d^{d-1} \mathbf{x}_{\parallel} \left\{ b_0 \left(\frac{\psi}{\beta} \right) \frac{1}{[\psi(\mathbf{x}_{\parallel})]^d} + b_2 \left(\frac{\psi}{\beta} \right) \frac{(\partial\psi)^2}{[\psi(\mathbf{x}_{\parallel})]^d} \right\}$$

Dimensionless functions b_0 and b_2 can be obtained from the knowledge of the Casimir free energy for small departures around the $\psi(\mathbf{x}_{\parallel}) = a = \text{constant}$ case.

$$\begin{aligned}
 b_0(\xi) &= \frac{\xi}{2} \sum_{n=-\infty}^{+\infty} \int \frac{d^{d-1} \mathbf{p}_{\parallel}}{(2\pi)^{d-1}} \log \left[1 - e^{-2\sqrt{(2\pi n \xi)^2 + \mathbf{p}_{\parallel}^2}} \right] \\
 &= \xi \frac{\pi^{(1-d)/2}}{2^{d-1} \Gamma(\frac{d-1}{2})} \sum_{n=-\infty}^{+\infty} \int_0^{\infty} d\rho \rho^{d-2} \log \left[1 - e^{-2\sqrt{(2\pi n \xi)^2 + \rho^2}} \right]
 \end{aligned}$$

$$b_2(\xi) = \frac{1}{2} \left[\frac{\partial F^{(2)}(\xi; n, |\mathbf{l}_{\parallel}|)}{\partial |\mathbf{l}_{\parallel}|^2} \right]_{n \rightarrow 0, |\mathbf{l}_{\parallel}| \rightarrow 0} .$$

where

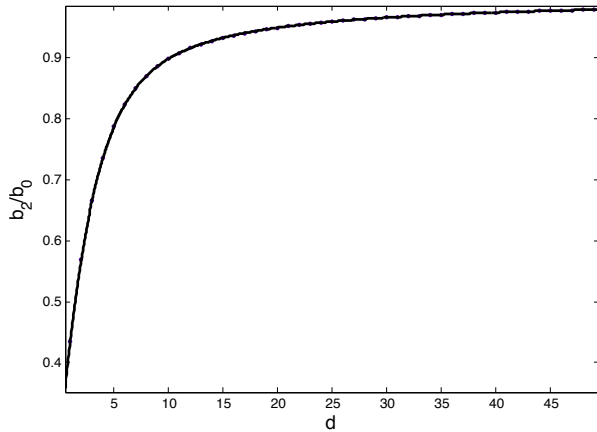
$$\begin{aligned}
 F^{(2)}(\xi; n, |\mathbf{l}_{\parallel}|) &= -2\xi \sum_{m=-\infty}^{+\infty} \int \frac{d^{d-1} \mathbf{p}_{\parallel}}{(2\pi)^{d-1}} \left\{ \right. \\
 &\quad \times \frac{\sqrt{(2\pi m \xi)^2 + \mathbf{p}_{\parallel}^2}}{1} \frac{\sqrt{(2\pi(m+n)\xi)^2 + (\mathbf{p}_{\parallel} + \mathbf{l}_{\parallel})^2}}{1} \\
 &\quad \left. \times \frac{1}{1 - \exp \left[-2\sqrt{(2\pi m \xi)^2 + \mathbf{p}_{\parallel}^2} \right]} \frac{1}{\exp \left\{ 2\sqrt{[2\pi(m+n)\xi]^2 + (\mathbf{p}_{\parallel} + \mathbf{l}_{\parallel})^2} \right\} - 1} \right\}
 \end{aligned}$$

Low and high temperature limits

Limits of the two coefficients b_0 and b_2 ; determine the form of the DE in the corresponding limits.

Relevant scale to compare the temperature with: inverse of the distance between the mirrors. Zero temperature limit: $\xi \rightarrow 0$

	$\frac{b_2(d)}{b_0(d)}$	\sim
$d = 1$	$\frac{1}{\pi^2} \left(1 + \frac{\pi^2}{3}\right)$	0.435
$d = 2$	$\frac{1+6\zeta(3)}{12\zeta(3)}$	0.569
$d = 3$	$2/3$	0.667
$d = 4$	$\frac{-\zeta(3)+10\zeta(5)}{12\zeta(5)}$	0.737
$d = 5$	$\frac{10\pi^2-21}{10\pi^2}$	0.787
$d = 6$	$\frac{-2\zeta(5)+7\zeta(7)}{6\zeta(7)}$	0.824



High temperature limit

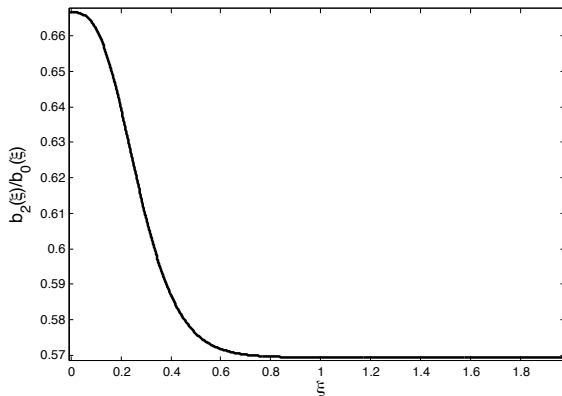
$$\xi \gg 1$$

$$[\Gamma_\beta(\psi, d)]_{\psi/\beta \gg 1} \sim \frac{1}{\beta} \int d^{d-1} \mathbf{x}_\parallel \left\{ b_0(d-1) \frac{1}{[\psi(\mathbf{x}_\parallel)]^{d-1}} + b_2(d-1) \frac{(\partial\psi)^2}{[\psi(\mathbf{x}_\parallel)]^{d-1}} \right\}.$$

(‘dimensional reduction’ phenomenon at high temperatures).
 In particular, for $d = 3$:

$$[\Gamma_\beta(\psi, 3)]_{\psi/\beta \gg 1} \sim -\frac{\zeta(3)}{16\pi\beta} \int d^2 \mathbf{x}_\parallel \frac{1}{[\psi(\mathbf{x}_\parallel)]^2} \{1 + 0.569 (\partial\psi)^2\}.$$

Ratio between $b_2(\xi)$ and $b_0(\xi)$ as a function of ξ ($d = 3$).



Neumann conditions

Now:

$$\Gamma_{\beta}(\psi) = \int d^{d-1} \mathbf{x}_{\parallel} \left\{ c_0 \left(\frac{\psi}{\beta} \right) \frac{1}{[\psi(\mathbf{x}_{\parallel})]^d} + c_2 \left(\frac{\psi}{\beta} \right) \frac{(\partial\psi)^2}{[\psi(\mathbf{x}_{\parallel})]^d} \right\}$$

with two (new) dimensionless functions c_0 and c_2 .

- An explicit calculation shows that the NTLO term is local and quadratic in derivatives for $d \neq 2$.
- $d = 2$ at zero temperature, or when $d = 3$ and the temperature is not zero: NTLO term is nonlocal.

The coefficients

The zero order term coincides with the one for the Dirichlet case;
 namely: $c_0 = b_0$.

	$\frac{c_2(d)}{c_0(d)}$	\sim
$d = 1$	$\frac{1}{3} - \frac{\zeta(0)}{3\zeta(2)}$	0.435
$d = 3$	$\frac{2}{3} - \frac{4\zeta(2)}{3\zeta(4)}$	-1.360
$d = 4$	$\frac{5}{6} - \frac{19\zeta(3)}{12\zeta(5)}$	-1.002
$d = 5$	$1 - \frac{9\zeta(4)}{5\zeta(6)}$	-0.915
$d = 6$	$\frac{7}{6} - \frac{2\zeta(5)}{\zeta(7)}$	-0.890
$d = 7$	$\frac{4}{3} - \frac{46\zeta(6)}{\zeta(8)}$	-0.886

Neumann conditions, $d = 2$

$d = 2$: not possible to compute the NTLO term coefficient by introducing the derivative with respect to $|l_{\parallel}|^2$ inside the integral \Rightarrow signal of branch cut at zero momentum. Also at $d = 3$ with $T > 0$ (thermal zero mode).

For small departures of the plane-plane geometry, up to second order in η ,

$$\Gamma_{\infty}^{(2)} = \frac{1}{2} \int \frac{d^2 k_{\parallel}}{(2\pi)^2} [g^{(2)}(k_{\parallel})]_{k_0 \rightarrow 0} |\tilde{\eta}(\mathbf{k}_{\parallel})|^2,$$

with:

$$g^{(2)}(k_{\parallel}) = -2 \int \frac{d^2 p_{\parallel}}{(2\pi)^2} \frac{[p_{\parallel} \cdot (p_{\parallel} + k_{\parallel})]^2}{|p_{\parallel}| |p_{\parallel} + k_{\parallel}|} \frac{1}{1 - e^{-2a|p_{\parallel}|}} \\ \times \frac{1}{e^{2a|p_{\parallel} + k_{\parallel}|} - 1}.$$

The form factor $g^{(2)}(k_{\parallel})$ does not admit an expansion in powers of k_{\parallel}^2 , for $d = 2$ (IR logarithmic divergence at $p_{\parallel} = 0$).

Nonlocal NTLO term

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Nonlocality related to the existence of massless modes in the theory. These modes are generally allowed by Neumann but not for Dirichlet boundary conditions (mass gap of order $1/a$).

Similar non-analyticity for $d = 3$ at finite temperature.

Higher orders

Next to NTLO terms:

$$|\partial\psi|^4, \psi|\partial\psi|^2\partial^2\psi, \psi^2\partial^2\psi\partial^2\psi$$

$$\psi^2\partial_\alpha\partial_\beta\psi\partial_\alpha\partial_\beta\psi, \psi^3\partial^2\partial^2\psi$$

(and terms more derivatives for higher orders).

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(and terms more derivatives for higher orders).

¿ Can the free energy be expanded in local terms up to any desired order? Necessary condition: the expansion must hold true for a particular case: when $\psi = a + \eta$, with $\eta \ll 1$, and one keeps just the quadratic term in η .

$T = 0$, $d = 3$ dimensions. The quadratic contribution to the Casimir energy:

$$\Gamma_\infty^{(2)} = \frac{1}{2a^5} \int \frac{d^2\mathbf{k}_\parallel}{(2\pi)^2} h^{(2)}(0, a\mathbf{k}_\parallel) |\tilde{\eta}(\mathbf{k}_\parallel)|^2,$$

where $h^{(2)}$ depends on the boundary conditions.

- On general grounds, we expect the form factor to be a function of $a(\omega^2 + \mathbf{k}_{\parallel}^2)^{1/2} \equiv a|k_{\parallel}|$, and explicit calculations confirm this fact.
- The Casimir energy will not admit an expansion in derivatives if the form factor includes odd powers or logs of its argument.
- At higher orders, and relaxing the condition of a quadratic approximation in η new non analytic terms may arise, which can be of the same order in the DE as the ones that come from the term quadratic in η .

Electromagnetic field

Extension of the improved PFA to the case of two imperfect thin mirrors.

F. Lombardo, F.D. Mazzitelli and C.D. F., Phys. Rev. D85 (2012) 125037

The action \mathcal{S} for this model is assumed to have the structure:

$$\begin{aligned}\mathcal{S} &= \mathcal{S}(A; y_L, y_R) \\ &= \mathcal{S}_0(A) + \mathcal{S}_L(A; y_L) + \mathcal{S}_R(A; y_R),\end{aligned}$$

where A denotes the 4-potential and \mathcal{S}_0 its free action:

$$\mathcal{S}_0(A) = \int d^4x \left[\frac{1}{4} F_{\mu\nu} F_{\mu\nu} + \frac{b}{2} (\partial_\mu A_\mu)^2 \right],$$

\mathcal{S}_L and \mathcal{S}_R are terms that couple A to each mirror, with y_L and y_R denoting parametrizations of their respective surfaces.

Assuming that \mathcal{S}_Σ is quadratic in A_μ :

$$\mathcal{S}_\Sigma(A; y) = \frac{1}{4} \int d^3\sigma \sqrt{g(\sigma)} d^3\sigma' \sqrt{g(\sigma')} F_{\alpha\beta}(\sigma) F_{\alpha'\beta'}(\sigma') \pi^{\alpha\beta\alpha'\beta'}(\sigma, \sigma'),$$

The model and its PFA

We shall consider a particular class of models, corresponding to:

$$\mathcal{S}_\Sigma(A; y) = \frac{\lambda}{4} \int d^3\sigma \sqrt{g(\sigma)} F_{\alpha\beta} F^{\alpha\beta},$$

where λ is a constant (can be extended to include frequency dependent couplings).

The improved PFA reads:

$$E_{\text{vac}} \simeq E_{\text{vac}}^{(0)} + E_{\text{vac}}^{(2)} + \dots$$

where the index denotes the order in derivatives.

$$E_{\text{vac}}^{(0)} = \int d^2\mathbf{x}_\parallel \int \frac{d^3k}{(2\pi)^3} \ln \left[1 - \frac{\left(\frac{|k|}{2}\right)^2}{\left(\frac{1}{\lambda_L} + \frac{|k|}{2}\right)\left(\frac{1}{\lambda_R} + \frac{|k|}{2}\right)} e^{-2|k|\psi(\mathbf{x}_\parallel)} \right].$$

NTLO term, imperfect mirrors

The NTLO correction to the PFA has the form

$$E_{\text{vac}}^{(2)} = \frac{1}{2} \int d^2 \mathbf{x}_{\parallel} c(\psi) (\partial_j \psi)^2.$$

Let us first consider the perfect-mirror limit. One can derive the idealized limit of perfect conductivity by taking $\lambda_1, \lambda_2 \rightarrow \infty$. The leading term reads, in this case

$$\left[E_{\text{vac}}^{(0)} \right]_{\text{perf}} = \int d^2 \mathbf{x}_{\parallel} \int \frac{d^3 k}{(2\pi)^3} \ln [1 - e^{-2|k| \psi(\mathbf{x}_{\parallel})}] = -\frac{\pi^2}{720} \int \frac{d^2 \mathbf{x}_{\parallel}}{\psi^3},$$

which is the well known PFA for perfect conductors.

NTLO for perfect mirrors ($\lambda \rightarrow \infty$) results in:

$$c = \frac{15 - \pi^2}{1080 \psi^3},$$

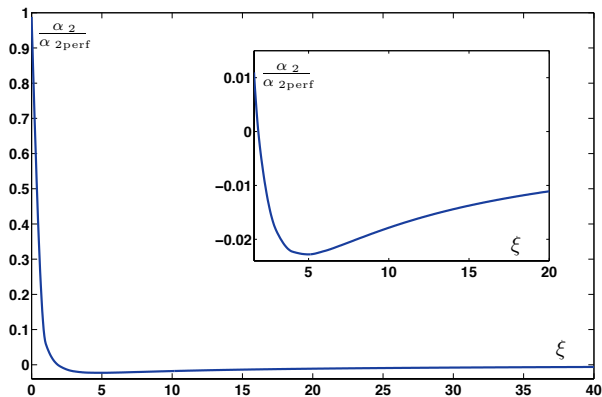
in agreement with the result obtained by Bimonte et al.

A model

$\lambda(k)$ proportional to $(k_0^2 + \mathbf{k}_{\parallel}^2)^{-1/2}$. Proportionality constant ξ .
By dimensional analysis,

$$E_{vac} \simeq \int d^2 \mathbf{x}_{\parallel} \frac{1}{\psi^3} [\alpha_1(\xi_L, \xi_R) + \alpha_2(\xi_L, \xi_R)(\partial_{\alpha} \psi)^2] .$$

Ratio between the coefficient of the NTLO correction $\alpha_2(\xi)$ and the corresponding value for the perfect case



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- DE useful in other contexts (we applied it to electrostatics).
- It has been generalized to imperfect mirrors, higher dimensions, and to finite temperature calculations.